

- (d) Discuss the open isotropic model of the universe given by the RW metric :

$$ds^2 = R^2(t) [-d\alpha^2 - \sin^2 \alpha (d\theta^2 + \sin^2 \theta d\phi^2)] + dt^2$$

with reference to dust and radiation distribution.

8

UNIT—V

9. Discuss the Schwarzschild interior solution as a simple stellar model. 16
10. Derive the TOV equation

$$p = \frac{-(\rho + p/c^2)(m + 4\pi k p r^3 / c^4)c^2}{r(r - 2m)} \text{ for}$$

the relativistic stellar structure.

16

AQ-1060

M.Sc. Part—II Semester—IV (CBCS) Examination MATHEMATICS

(General Relativity and Cosmology—II)

Paper—XVIII-(403)

Time : Three Hours]

[Maximum Marks : 80

Note :—Solve FIVE questions selecting one from each unit.

UNIT—I

1. (a) By considering the static homogeneous and spherically symmetric universe, obtain the Einstein's model of the universe. 8
- (b) Prove that there is no Doppler shift in the Einstein's universe. 8
2. (c) Prove that the geometry of the de-Sitter universe is that of the surface of a sphere embedded in five dimensional Euclidean space. 8
- (d) Discuss the velocity and acceleration of a particle in the de-Sitter universe and show that the particle at rest at origin will remain permanently at rest at the origin. 8

UNIT—II

3. (a) State and explain :
Cosmological principle. 4
- (b) By making the necessary assumptions, derive the Robertson-Walker line element. 12

4. (c) Prove that the quantity $\frac{1}{a^2}$ in the line element

$$ds^2 = -e^{g(t)} \left[\frac{dr^2}{1-r^2/a^2} + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2 \right] + dt^2$$
 is

the Riemannian curvature of the spatial extent at any given time t . 8

- (d) Considering the spatial part of the closed form of RW model, Prove that the total proper spatial volume at any given time t is $V = 2\pi^2 R^3$. 8

UNIT—III

5. (a) Discuss the motion of particle in R- W model and prove that a material particle does not experience the gravitational force. 8
- (b) Prove that the equation of light trajectory in the RW universe is given by

$$r = za \tan \left(\frac{e^{-bt_1} - e^{-bt_2}}{2ab} \right). \quad 8$$

6. (c) Prove that in RW model, there exists Red shift for expanding universe and a blue shift for contracting universe. 8
- (d) Prove that the equation for red shift z in terms of l is :

$$z = R'_0 \ell + \frac{R_0'^2}{2} (1 + q_0) \ell^2 + O(\ell^3),$$

where q_0 is the deceleration parameter. 8

UNIT—IV

7. (a) Obtain the fundamental equation of dynamic cosmology :

$$\dot{R}^2 + k = \frac{8\pi\rho R^2}{3}. \quad 8$$

- (b) Prove that the matter dominated era of the universe is governed by the equation :

$$\left(\frac{\dot{R}}{R_0} \right)^2 = H_0^2 \left\{ 1 - 2q_0 + 2q_0 \frac{R_0}{R} \right\},$$

where the symbols have their conventional meanings. 8

8. (c) Derive the equation for the age of the universe to, for Friedmann closed model in the form

$$t_0 = \frac{q_0}{H_0 (2q_0 - 1)^{3/2}} \left\{ \cos^{-1} \left(\frac{1}{q_0} - 1 \right) - \frac{1}{q_0} (2q_0 - 1)^{1/2} \right\}.$$

8